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Abstract A new model for the giant H II loops in IM and Sc galaxies is proposed. The ram pressure of a massive, collapsing gas cloud serves to confine the stellar
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propose that thebeating of several non-radial pulsation modes triggers transientlarge-scale density perturbations in a confined stellar wind stellar wind and
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The bubble in NGC 7635 is the result of a fast stellar wind expanding into the away from the cloud wall is confined by the ram pressure of the wind.
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Stellarator definition, an experimental plasma-physics device in which magnetic fields confine the plasma within a stellar evolution; stellar wind; stellarator;
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to determine the optimal shapes of the coils that created the magnetic field to confine the stellar parallax; stellar wind; stellarator; Stellaria; Stellaria

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Physical mechanisms for the bipolar jets around

Physical mechanisms for the bipolar jets around young stellar objects we see that 80 VUE & Winnewisser in order to confine a wind with $A_w = 4 \times 10^9$ Myr

The effect of stellar winds on the formation of a

is likely in the highly inhomogeneous gas distributions existing in molecular clouds that some stellar wind exits through a hole wind material may confine

Similarity solution for the interaction of the

the stellar wind, in the simple model the swept mass of interstellar medium by stellar wind is confined in the thin shell located at R_2 . The same

Stagnation and infall of dense clumps in the

all the while being confined by the ram pressure of the outflowing ambient wind. Stagnation and Infall of Dense Clumps in the Stellar Wind of Scorpii.

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with emphasis on the central competition between confinement by the magnetic field vs. expansion of the stellar wind can confine parts of the wind up

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characters is unusual because evolved hot stars are expected to rotate relatively slowly after braking by a strong stellar wind, which would confine other

Stellar wind information archive

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The magnetic field and wind confinement of

this star a consistent model of the large-scale magnetic field and of the associated magnetically confined wind and circumstellar stellar-wind streams from

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The velocities of the clumps relative to the ambient stellar wind can approach 2000 km s⁻¹, all the while being confined by the ram pressure of the outflowing

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